

ABUNDANCE OF C⁺/O⁺ INNER SOURCE PICKUP IONS PRODUCED BY SOLAR WIND RECYCLING, **NEUTRALIZATION, AND SPUTTERING**

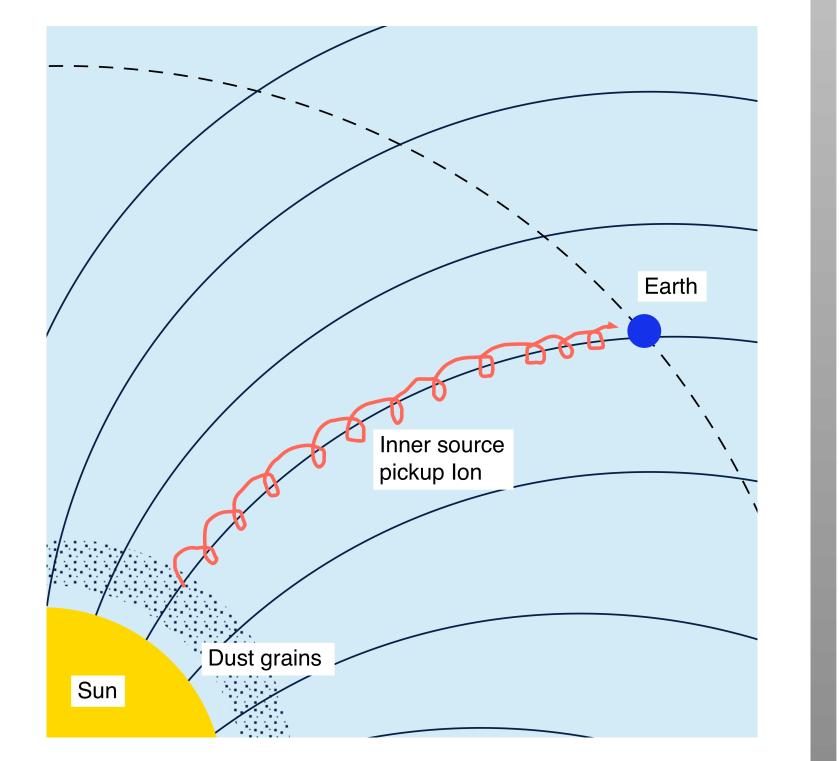
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MOTIVATION

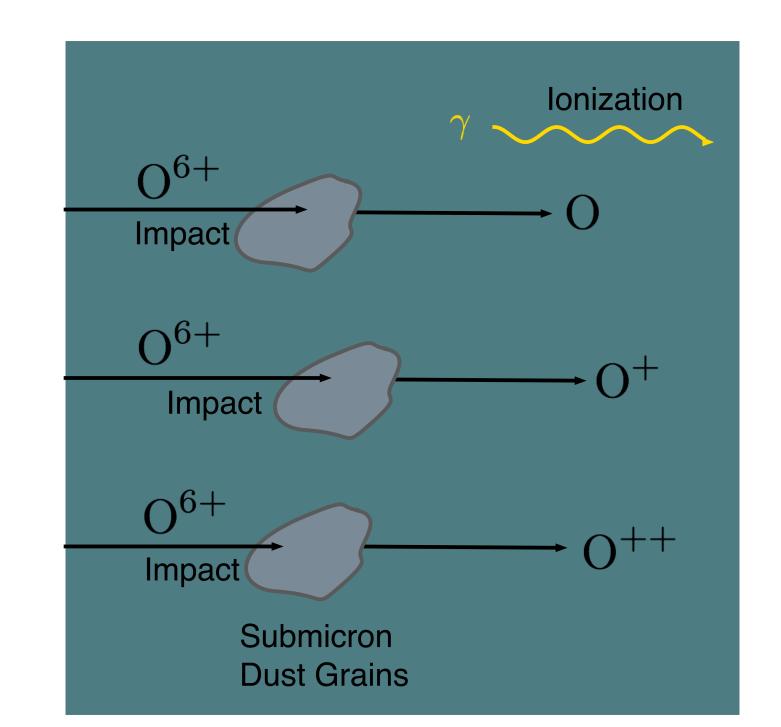
- Inner source pickup ions are singly charged ions that are produced near the sun.
- The mechanism behind the production of inner source pickup ions is not well understood
- Possible production mechanisms include:
- Solar wind recycling absorption and desorption of solar wind ions in dust grains
- Solar wind neutralization charge exchange through carbon foil-like dust grain population
- **Sputtering** from grains
- Dust-dust collisions
- Sun-grazing comets



SOLAR WIND NEUTRALIZATION

MECHANICAL PROCESS

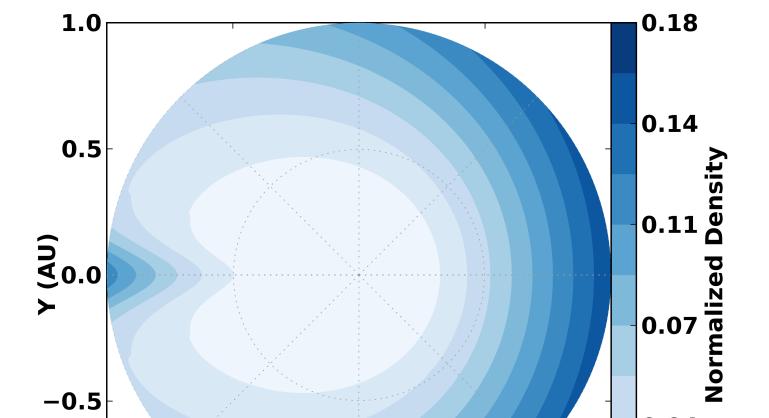
- Solar wind ions completely penetrate submicron sized dust grains
- Charge exchange leaves a fraction of neutral atoms, singly-charge ions, doubly-charged ions, and so on.
- The neutral atoms become singly-charged ions, and the singlycharge ions becomes doubly-charged ions once ionized by photoionization, charge exchange with solar wind protons, or electron impact.



INTERSTELLAR SOURCE

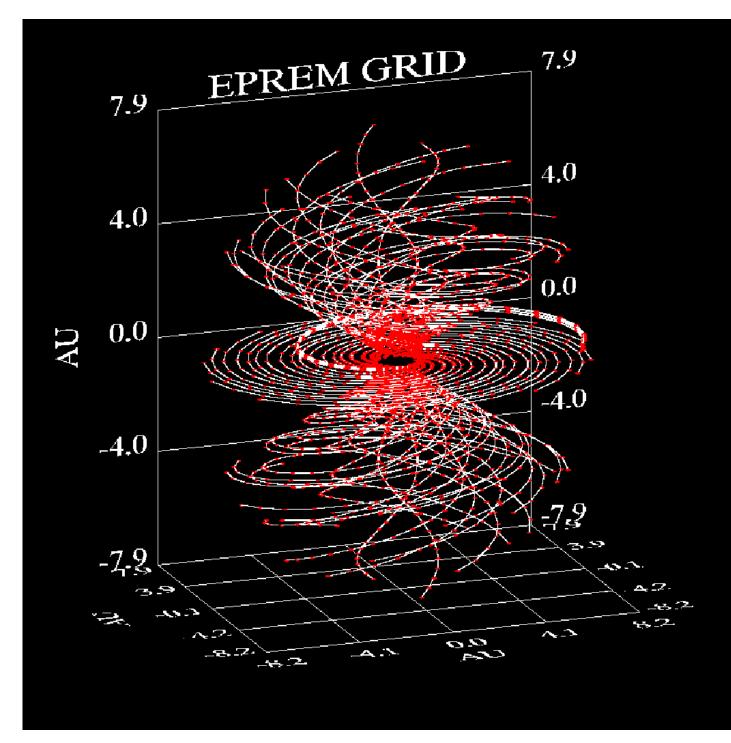
MECHANICAL PROCESS

- Interstellar neutral atoms penetrate the heliosphere and follow hyperbolic orbits due to the sun's gravity.
- The neutral atoms becomes ionized due to photoionization, charge exchange, and electron impact.
- Once ionized, they are picked up by the solar wind.



EPREM

- **Energetic Particle Radiation Environment Module (EPREM)**
- Models particle transport throughout the heliosphere.
- Solves the focused transport equation and convection-diffusion equation.
- Takes into account: convection, streaming, adiabatic cooling, adiabatic focusing, pitch-angle scattering, perpendicular diffusion, and drift.



Question How does C^+/O^+ produced by solar wind recycling, neutralization, and sputtering compare to observations?

Question How does the flux of C^+ and O^+ at 1 AU compare to observations?

Question Is there a dominant mechanism?

SOLAR WIND RECYCLING

MECHANICAL PROCESS

- Solar wind ions penetrate about 10-30 nm per keV and nucleon into the outer mineral layer of dust grains.
- Electron capture neutralizes the solar wind ion within the dust grain.
- Dust grains become saturated with neutralized solar wind atoms.
- Due to saturation, the solar wind atoms diffuse to the surface and desorb from the dust grain.

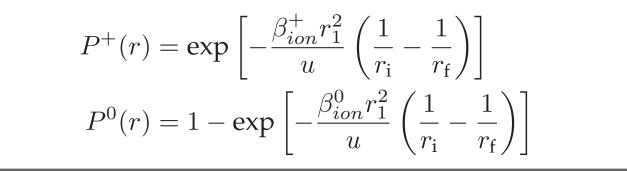
FORMALISM

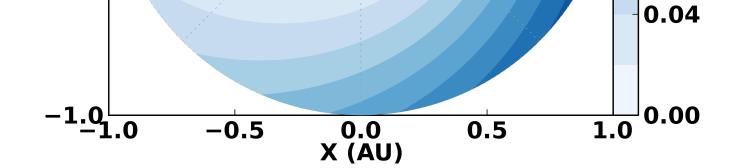
• Production of singly-charged pickup ions depends on the impact rate, charge state fraction after penetration, and the loss rate probability due to ionization.

• Neutral atoms produced after grain penetration have a probability of becoming singly-charged pickup ions due to ionization adding to the singly-charge pickup ions population.

$$S_{\text{neut}}^+(r) = \eta^+ S_{\text{impact}}(r) P^+(r) + \eta^0 S_{\text{impact}}(r) P^0(r)$$

• The survival and production probabilities beyond penetrating the grains depend on the ionization rates, distance traveled, and solar wind speed.





FORMALISM

- The hot gas model calculates the spatial distribution of interstellar neutral atoms within the heliosphere.
- The production rate per unit volume depends on the ionization rates and neutral density.

$$S_{\rm int}^+(r) = \left(\beta_{\rm ph} + \beta_{\rm ch} + \beta_{\rm el}\right) \left(\frac{r_1}{r}\right)^2 n_{\rm O}(r,\theta).$$

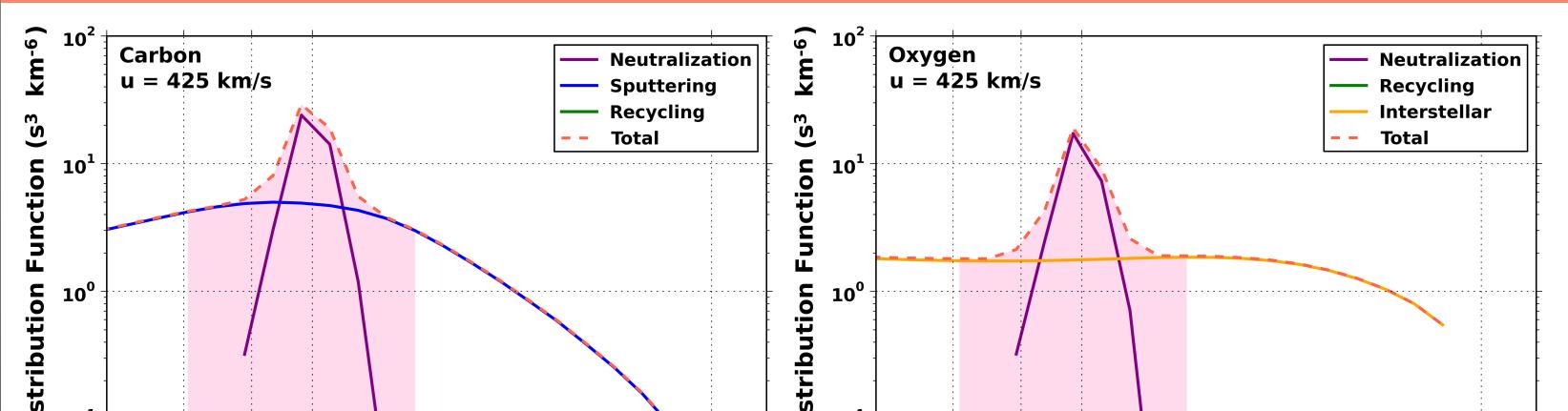
• Although electron impact scales slightly steeper than r^{-2} , we assume it scales as r^{-2} for simplicity.

RESULTS

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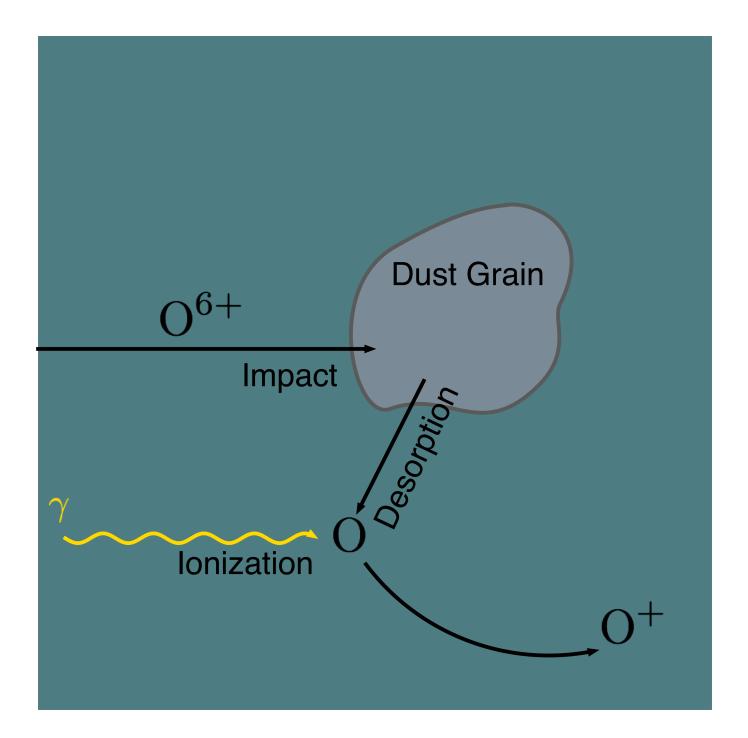


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PROCEDURE

- Add the production rate per unit volume of inner source and interstellar carbon and oxygen into EPREM.
- Simulate the transport of carbon and oxygen inside 1 AU.
- Integrate the velocity distribution function between $0.8 \le w \le 1.2$ to get the flux and abundance ratios.
- Compare to CTOF data.

- Once desorbed, the neutrals are ionized by photoionization, charge exchange, or electron impact.
- Due to the very close proximity to the sun, ionization occurs almost immediately.
- After ionization, the ions are picked up and convected out with the solar wind.



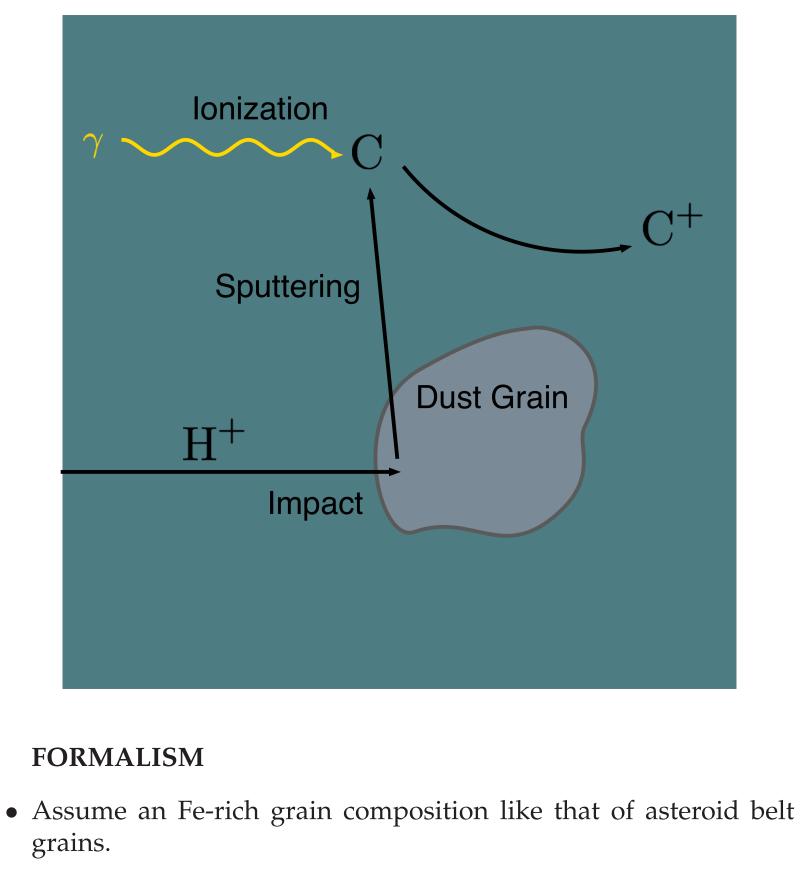
FORMALISM

• Impact rate per unit volume depends on the flux of solar wind heavy ions and cross section of dust grains.

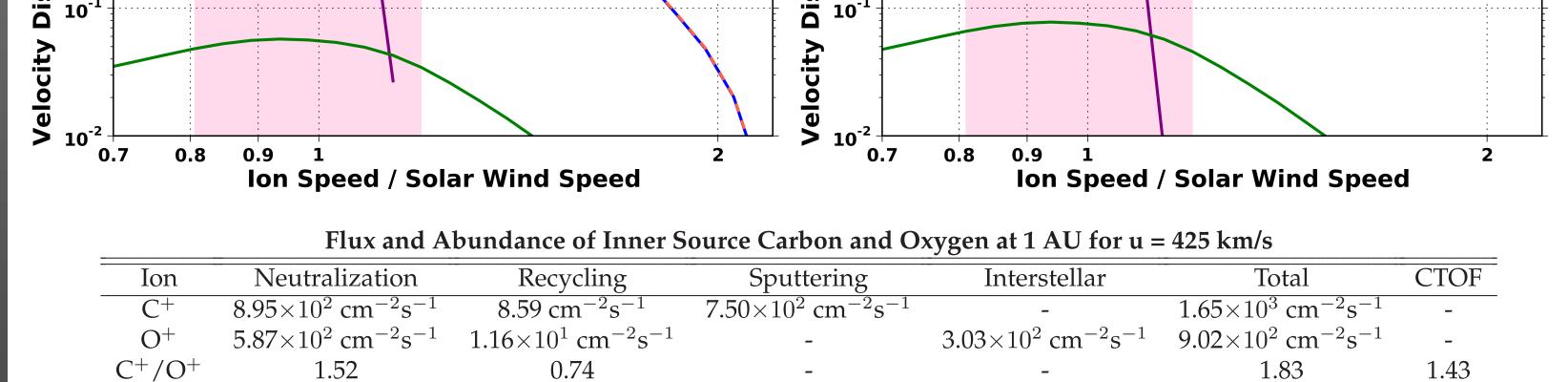
SPUTTERING

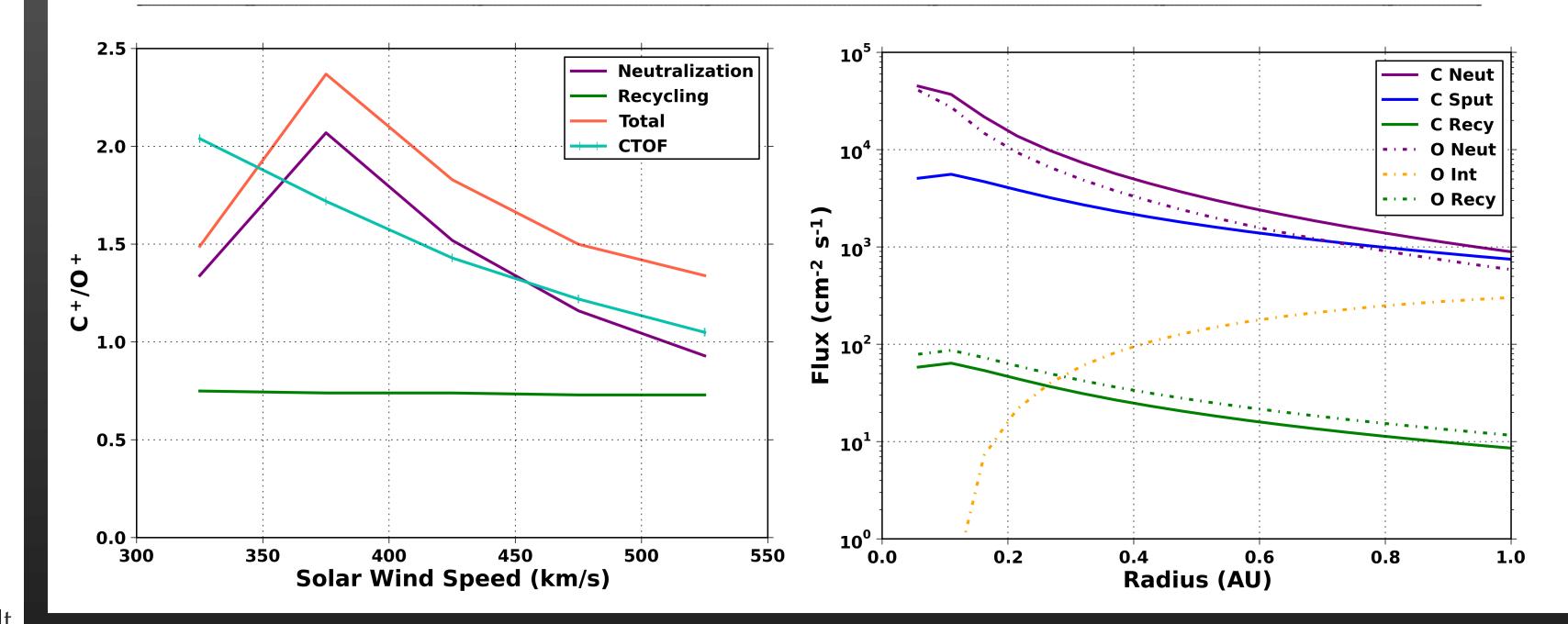
MECHANICAL PROCESS

- Solar wind protons bombard dust grains.
- Carbon is ejected from the grain when given an energy greater than the surface binding energy of the grain.



• Assume these Fe-rich grains make up the entire grain population.





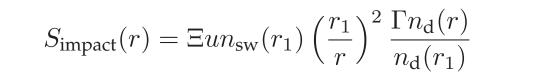
SUMMARY

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• The total C^+/O^+ ratio is comparable to CTOF for varying solar wind speeds.

REMAINING QUESTIONS

• How does the velocity distribution of C⁺ and O⁺ measured by



(1)

(2)

• Absorption, desorption, and ionization are assumed to be in equilibrium. Therefore the production rate of pickup ions equals the impact rate.

 $S_{\text{recv}}^+(r) = S_{\text{impact}}(r)$

• The rate of sputtered carbon is related to the rate of impacting protons by the sputtering yield.

 $S_{\text{sput}}^+(r) = YS_{\text{impact}}(r)$

• Sputtering yield of carbon based on laboratory experiments.

• Assume that the sputtering of carbon is proportional to the percent of carbon within the grain.

• The sputtering of oxygen within Fe-rich grains is negligible.

• The total C^+/O^+ ratio mainly follows solar wind neutralization. produced by solar wind recycling follows solar wind • C^+/O^+ C/O.

• The majority of inner source pickup ions produced near the sun are from solar wind neutralization, followed by sputtering.

• C^+ and O^+ produced by solar wind recycling are negligible. However, sputtering may still release solar wind C and O imbedded in the surface of the grain.

CTOF compare to the simulation results?

• What is the flux of C⁺ and O⁺ measured by CTOF?

• What causes C^+/O^+ to decrease for very slow solar wind speeds?

• How does each production mechanism depend on grain size?

• What is the distribution of each composition of grains near the sun, as compared to asteroid, meteoritic, and interstellar grains?

• Are implanted solar wind ions sputtered?