

In situ observations of EMIC waves and associated plasma properties in the inner magnetosphere: Comparisons of two events

J.-C. Zhang¹(jichun.zhang@unh.edu), L. M. Kistler¹, C. G. Mouikis¹, R. C. Allen¹, B. Klecker², J.-A. Sauvaud³, and M. W. Dunlop⁴

¹Space Science Center, University of New Hampshire, Durham, NH 03824, USA

²Max-Planck-Institut für extraterrestrische Physik, Garching, Germany ³CESR/CNRS, Avenue du Colonel Roche, 31028 Toulouse Cédex 4, France

⁴Space Sciences Division, SSTD, Rutherford Appleton Laboratory, Chilton, DIDCOT, Oxfordshire, OX11 0QX, United Kingdom

2011 CEDAR-GEM Joint Summer Workshop, Santa Fe, NM, June 26 – July 1, 2011



Paper#: 94



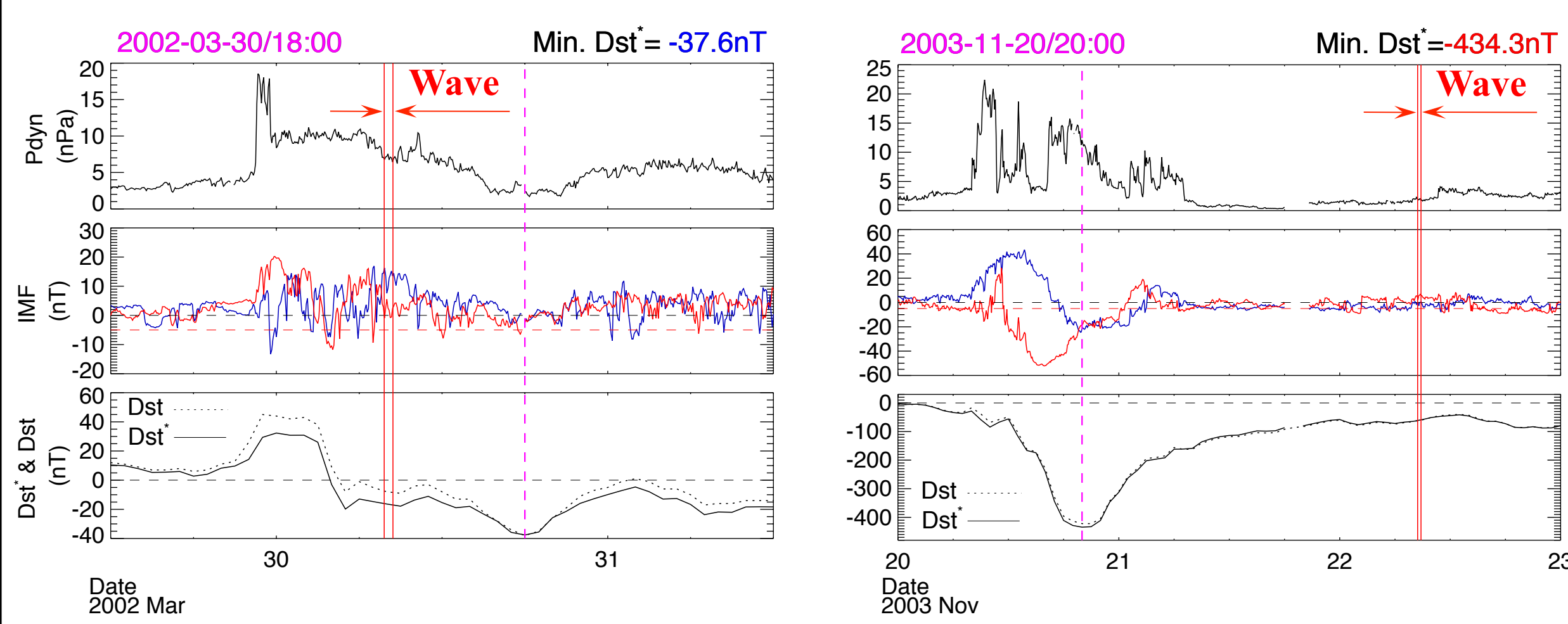
Abstract

Electromagnetic ion cyclotron (EMIC) waves, which are normally induced by anisotropic distributions of energetic H⁺, play an important role in the overall dynamics of the Earth's magnetosphere, contributing to the loss and energization of particles. In this study, two EMIC wave events and associated plasma properties are examined and compared against each other. The 30 Mar. 2002 event occurred in the main phase of a mild storm, while the 22 Nov. 2003 event took place in the recovery phase of a rare super-storm. During both events, *in-situ* observed plasma conditions were favorable to the occurrence of the EMIC instability and also indicated the resonant interactions of particles with the waves. However, in the former event they are consistent with linear theory but in the latter one they are not. The possible reason is that in one case Cluster detected the wave near its source region but in the other case the observed wave had propagated for a while and already experienced a polarization reversal.

Motivation

1. To investigate the causes, effects, and differences of the two EMIC wave events observed by Cluster in the inner magnetosphere, which could cause pitch-angle scattering and loss of relativistic e⁻.
2. To extend the case and statistical studies of EMIC wave-associated He⁺ heating events in the outer magnetosphere [Zhang *et al.*, 2010; 2011].
3. To follow up the case studies of the 30 Mar. 2002 wave [Pickett *et al.*, 2010] and the 22 Nov. 2003 wave [Engebretson *et al.*, 2007].

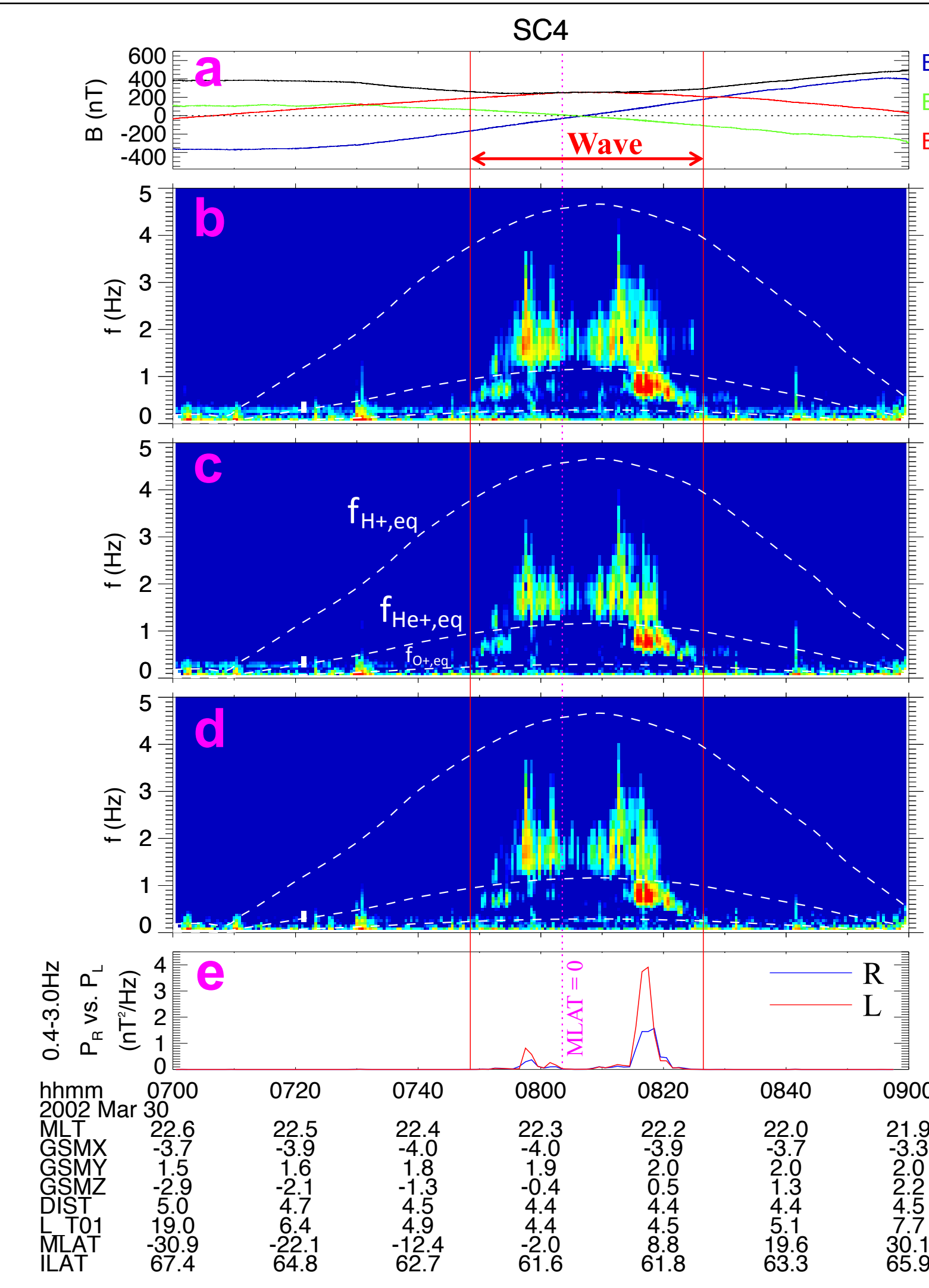
Solar Wind/Geomagnetic Condition



- ❖ **Wave on 30 Mar. 2002 [left panel]:** Observed in the main phase of a weak storm. During the wave period,
 - Avg. $P_{dyn} = 7.0$ nPa
 - Avg. IMF $B_y = 12.8$ nT
 - Avg. IMF $B_z = 1.6$ nT
 - Avg. Dst^* (pressure-corrected Dst) = -16.3 nT

- ❖ **Wave on 22 Nov. 2003 [right panel]:** Observed in the recovery phase of a super-storm. During the wave period,
 - Avg. $P_{dyn} = 1.8$ nPa
 - Avg. IMF $B_y = -1.8$ nT
 - Avg. IMF $B_z = 3.5$ nT
 - Avg. $Dst^* = -61.1$ nT

Magnetic Field & Wavelet Analysis

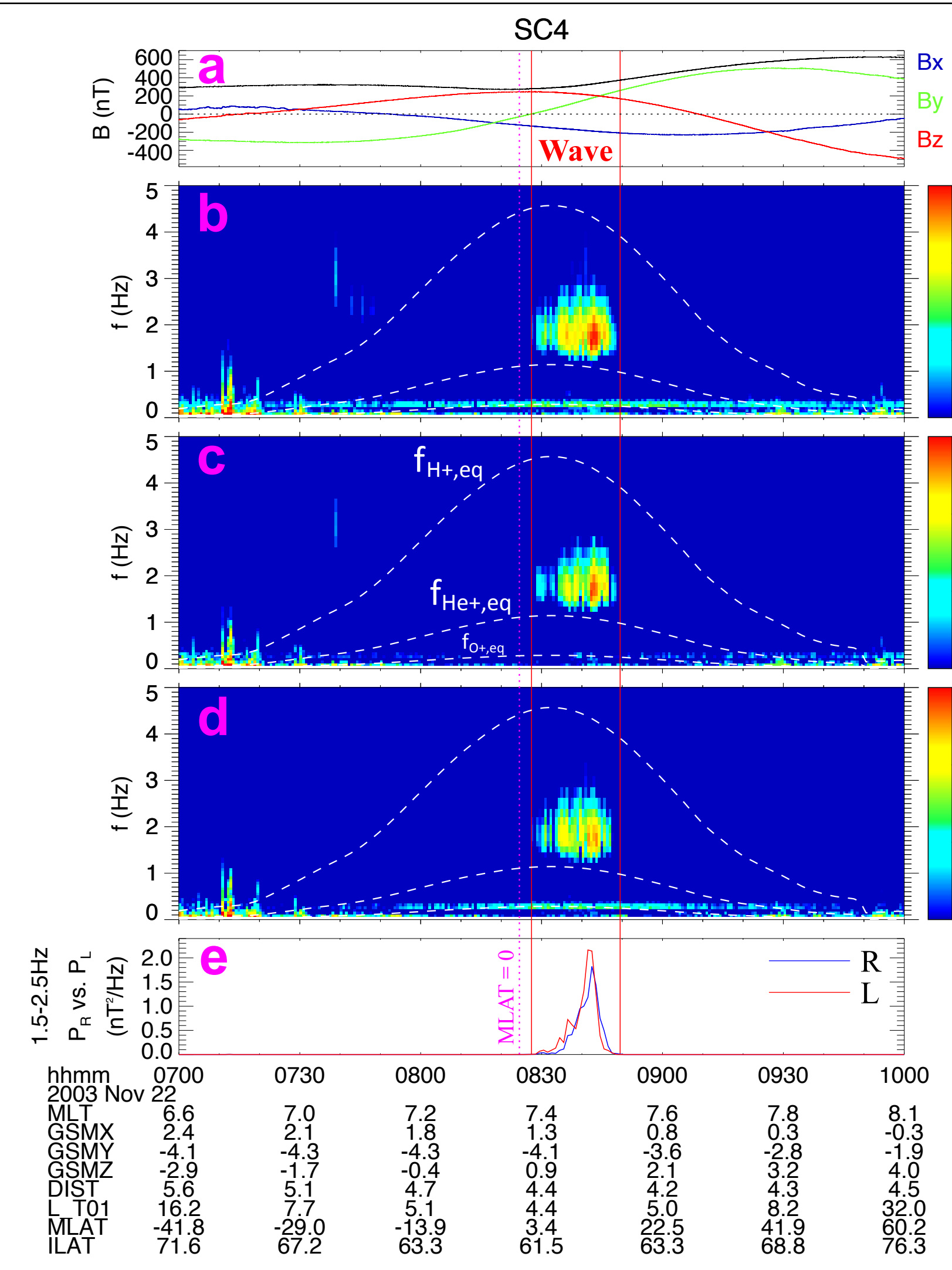


Total wave power

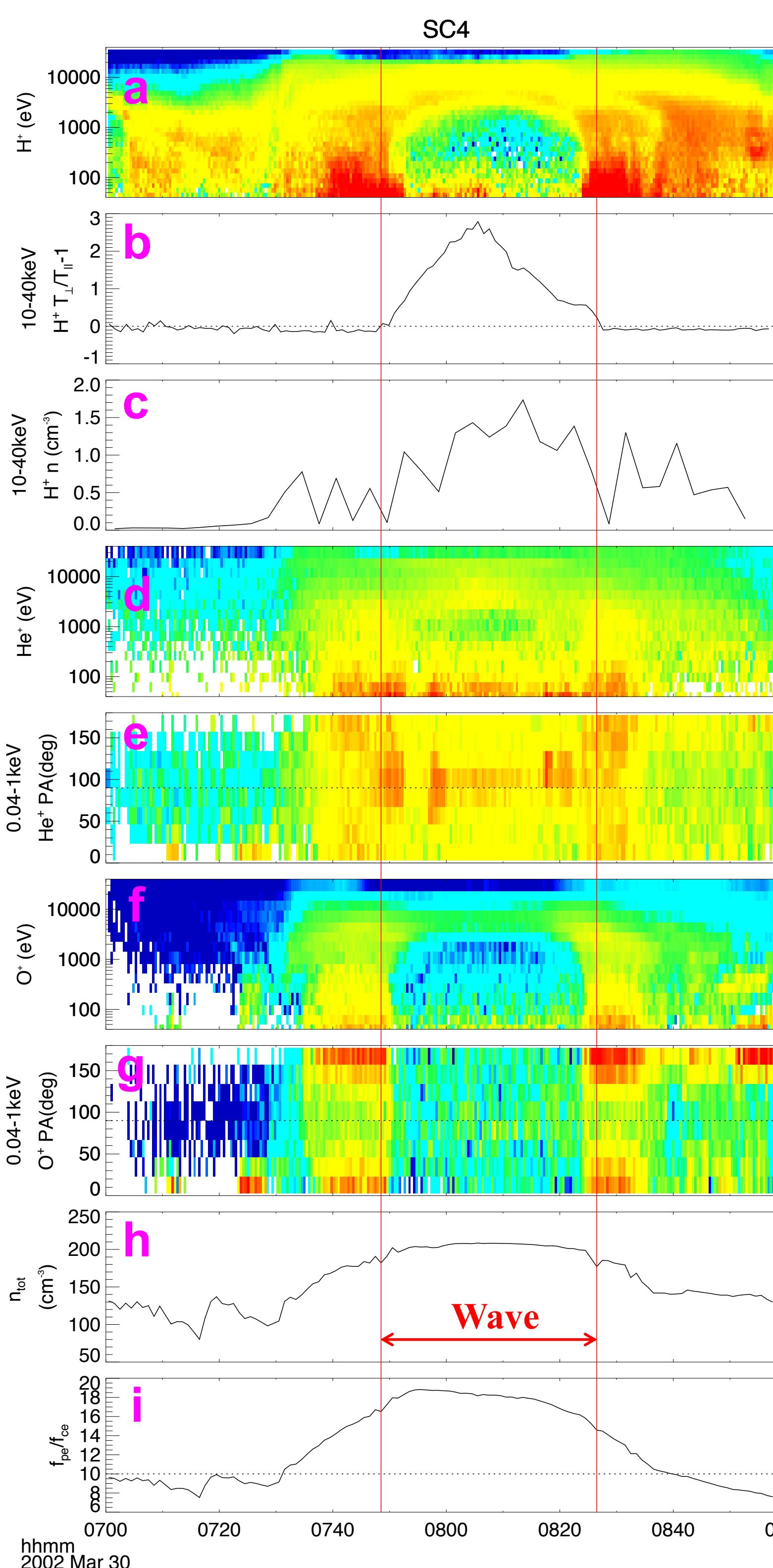
Right-hand polarized wave power

Left-hand polarized wave power

Average wave power in the frequency range of wave activity



Plasma Data



H⁺ spectrograms

Hot H⁺ anisotropy

Hot H⁺ density

He⁺ spectrograms

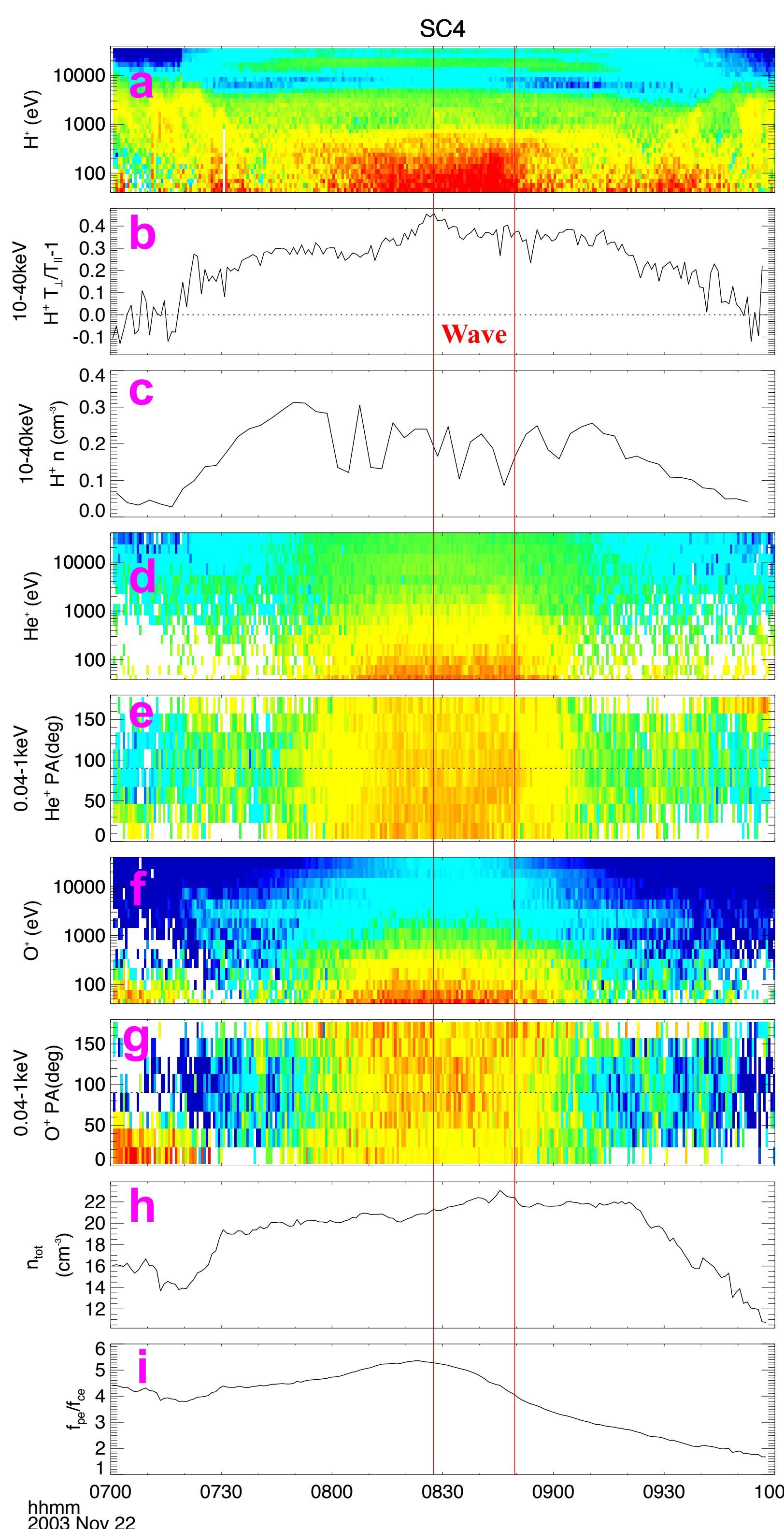
Cool He⁺ pitch angle distribution

O⁺ spectrograms

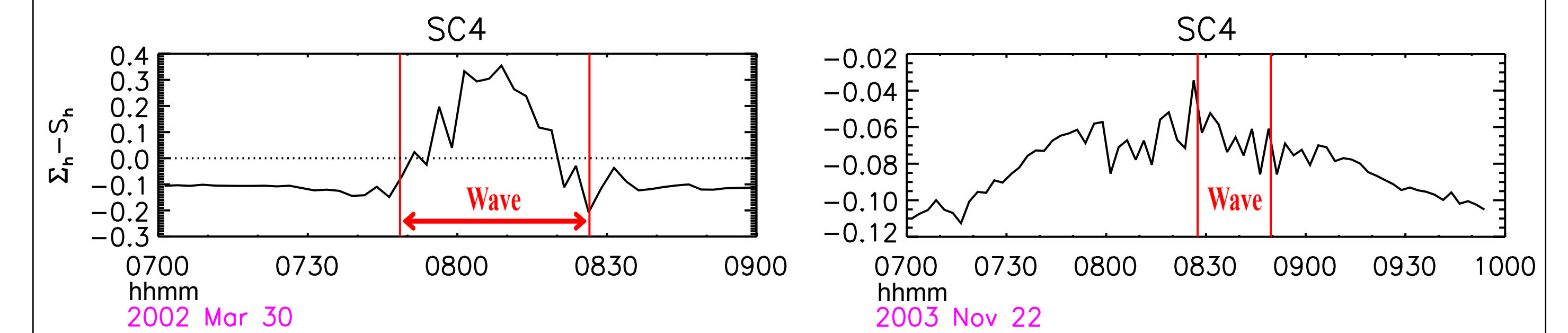
Cool O⁺ pitch angle distribution

Total ion density ($= n_e$)

e⁻ plasma/gyrofrequency ratio



Comparison with Linear Theory



❖ **Linear Theory:** EMIC wave activity can occur only if $\Sigma_h - S_h > 0$ [Gary *et al.*, 1994], where Σ_h is the observational EMIC growth parameter and S_h is the EMIC instability threshold.

❖ $\Sigma_h - S_h > 0$ around the center of the event on 30 Mar. 2002 [left panel]. However, $\Sigma_h - S_h < 0$ during the 22 Nov. 2003 event [right panel].

Summary & Discussion

	30 Mar. 2002	22 Nov. 2003
Geomagnetic Condition	Main phase of a small storm	Recovery phase of a super-storm
Location	L~4.4, MLT~22.3, MLAT~0°	L~4.4, MLT~7.4, MLAT~3.4°
Region	Equatorial plasmapause	Above the equatorial plane
Wave Frequency	Around $f_{He^{+},eq}$	Below $f_{He^{+},eq}$
Total Wave Power	up to 50 nT ² /Hz	up to 8 nT ² /Hz
Wave Polarization	Left-handed	Nearly linear
Hot H⁺ Anisotropy	up to ~3	up to ~0.5
Hot H⁺ Density	up to 1.7 cm ⁻³	up to 0.3 cm ⁻³
He⁺ Heating	Strongly	Mildly
O⁺ Heating	Slightly	N/A due to RB e ⁻ contamination
n_{tot}	up to 200	up to 23
f_{pe}/f_{ce}	up to 18 (> 10)	up to 5 (< 10)
Theory Consistency	Yes	No
Explanation	Newly generated Close to the wave source region	Propagated & polarization reversed Away from the wave source region

● **Future Work:** 1) create lists of all Cluster-observed EMIC waves & associated plasma properties; 2) understand discrepancies between observations & wave theories/modeling.

References

Gary, S. P., et al. (1994), *JGR*, 99(A12).
Pickett, J. S., et al. (2010), *GRL*, 37(L09104).
Engebretson, M. J., et al. (2007), *PPS*, 55
Zhang, J.-C., et al. (2010), *JGR*, 115(A06212).
Zhang, J.-C., et al. (2011), *JGR*, under review.

Acknowledgements

Work at UNH was supported by NASA through grant NNX11AB65G and by the RBSP-ECT project. The authors thank the NSSDC OMNI Data Center and the Cluster team for making their data and/or software available.