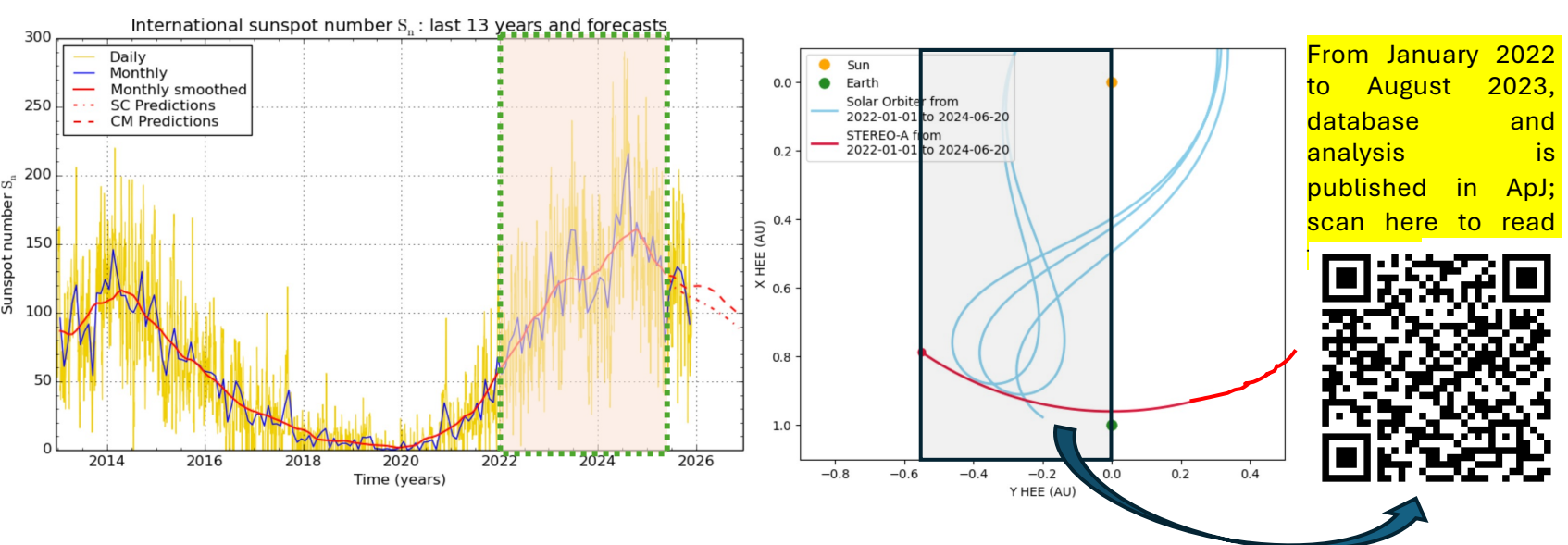


## 1. Introduction

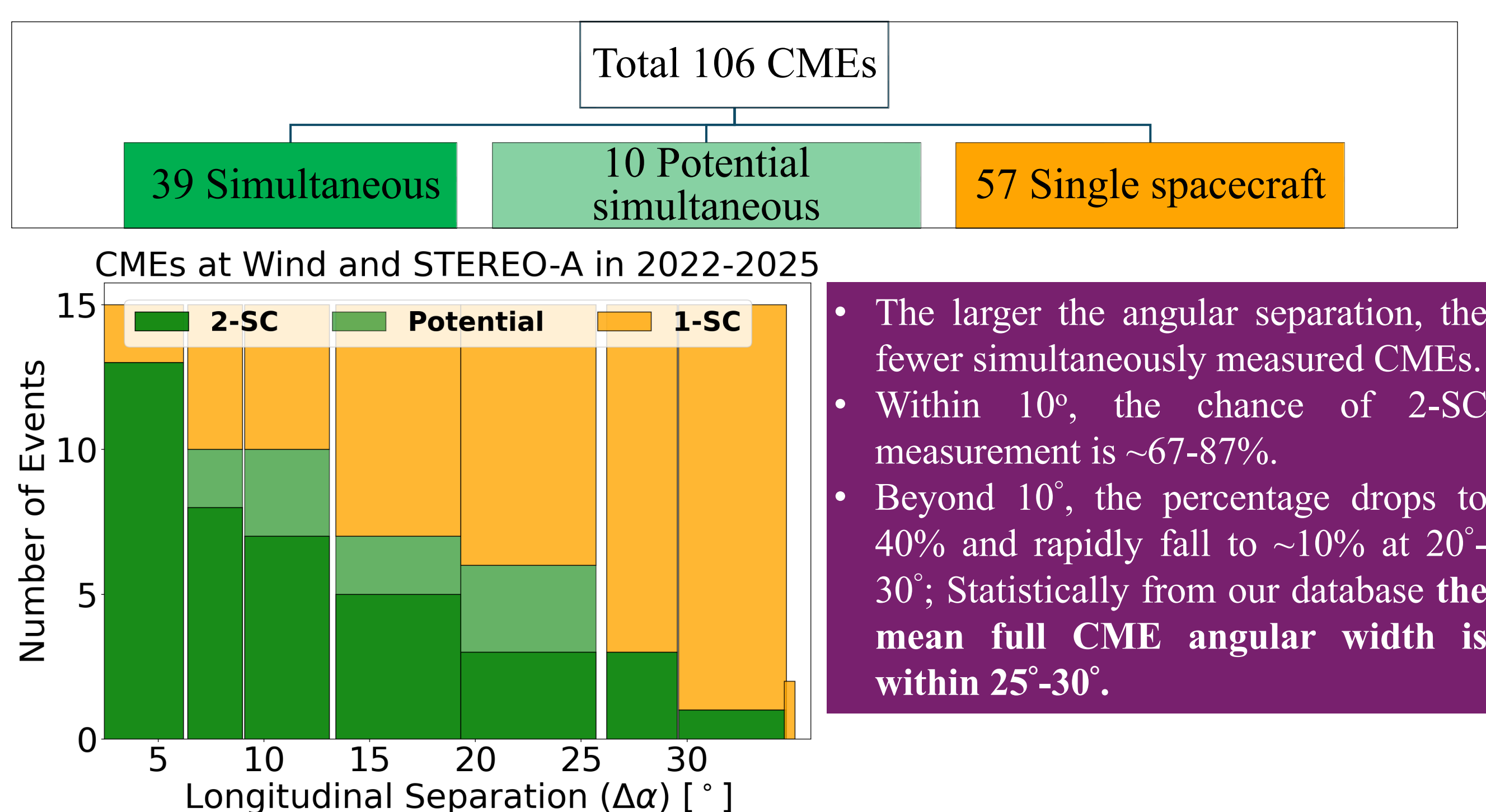


Between January 2022 and May 2025, STEREO-A moved from  $35^\circ$  east of Earth to west.

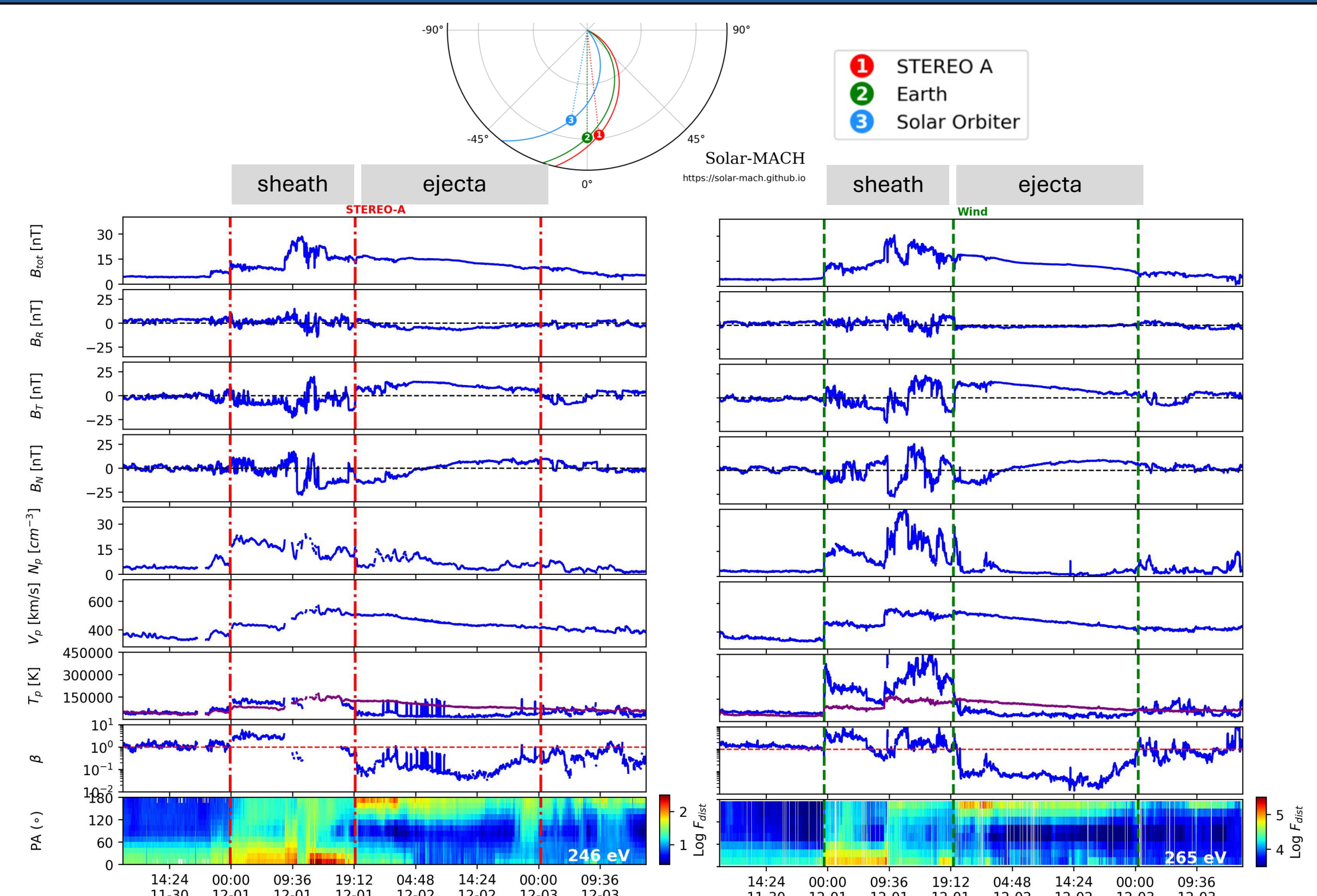
Objectives:

1. Analyze *longitudinal size and various properties of CMEs* using multipoint observations from STEREO-A and L1 spacecraft.
2. Perform detailed assessments of *CME-driven shock and ejecta morphology* and their propagation characteristics.

## 2. Database



## 3. CME event on 2023-12-01



## 7. Conclusions

- Built a catalog of 106 CMEs combined by Wind and STEREO-A within 2022 to May 2025. Only 49 CMEs (~46%) measured by 2-SC while their angular separation ranges between  $35^\circ$  each side from the Sun-Earth line.
- From  $0^\circ$  and  $35^\circ$  separation, percentage of detecting same CME simultaneously by 2-SC reduced 85% to below 10%.

*Average CME full angular width near 1 au is likely to be  $25^\circ$  to  $30^\circ$ , smaller than previously thought.*

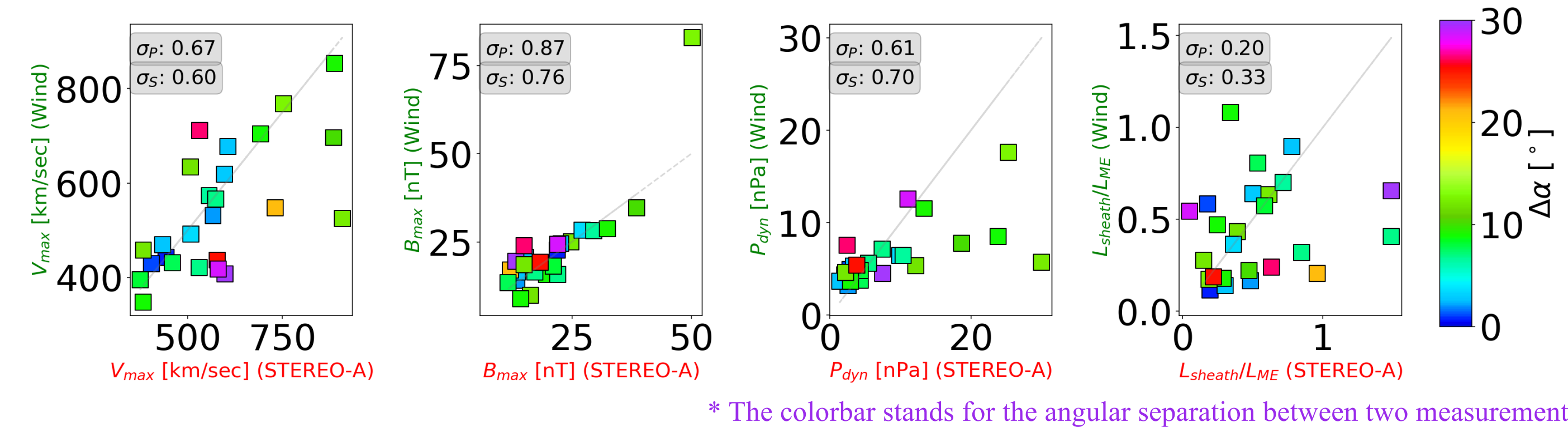
- Average properties of sheath and ME such as, speed, magnetic field are found to be consistent, regardless of separation between measurements.
- However, close inspection by using RMSE, correlation coefficients suggests that locally, properties vary significantly within such a smaller longitudinal width of CMEs.

*MEs are globally coherent structure within the scale of  $10^\circ$  to  $20^\circ$ , however, field components do vary significantly in smaller scale.*

- Globally, shocks are consistent with a circular cross-section.

## 4. Average properties of Sheath and Magnetic Ejecta (ME)

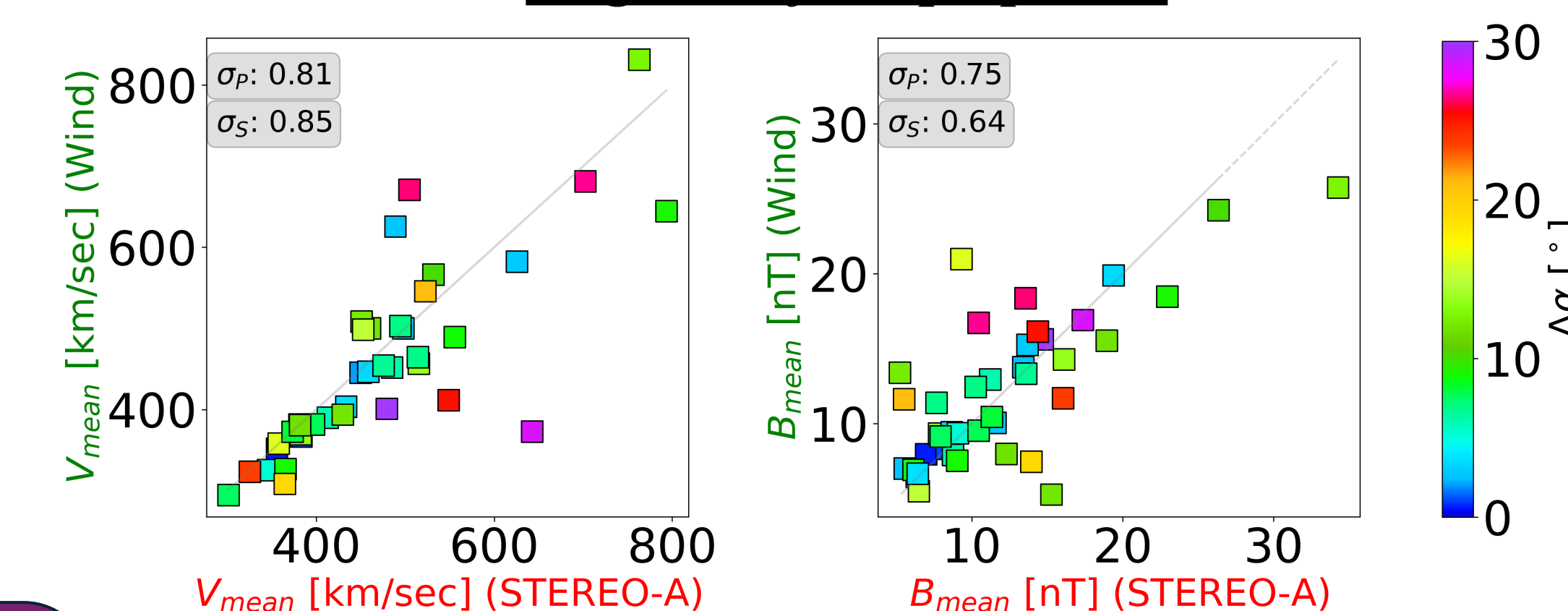
### Sheath properties



- 25 of 39 simultaneous events show sheath signatures at both spacecraft.
- Peak magnetic field is highly consistent, while peak speed and dynamic pressure show moderate correlation; the sheath-to-ejecta length ratio is poorly correlated because each spacecraft crosses different parts of the CME.
- Mean magnetic field and mean dynamic pressure are strongly consistent, whereas sheath length and average sheath speed are only moderately correlated, with substantial variation even for separations  $<10^\circ$ .

- While a faster ME speed (as seen at STEREO-A) is generally expected to correspond to a smaller sheath-to-ME length ratio, L1 measurements doesn't align.

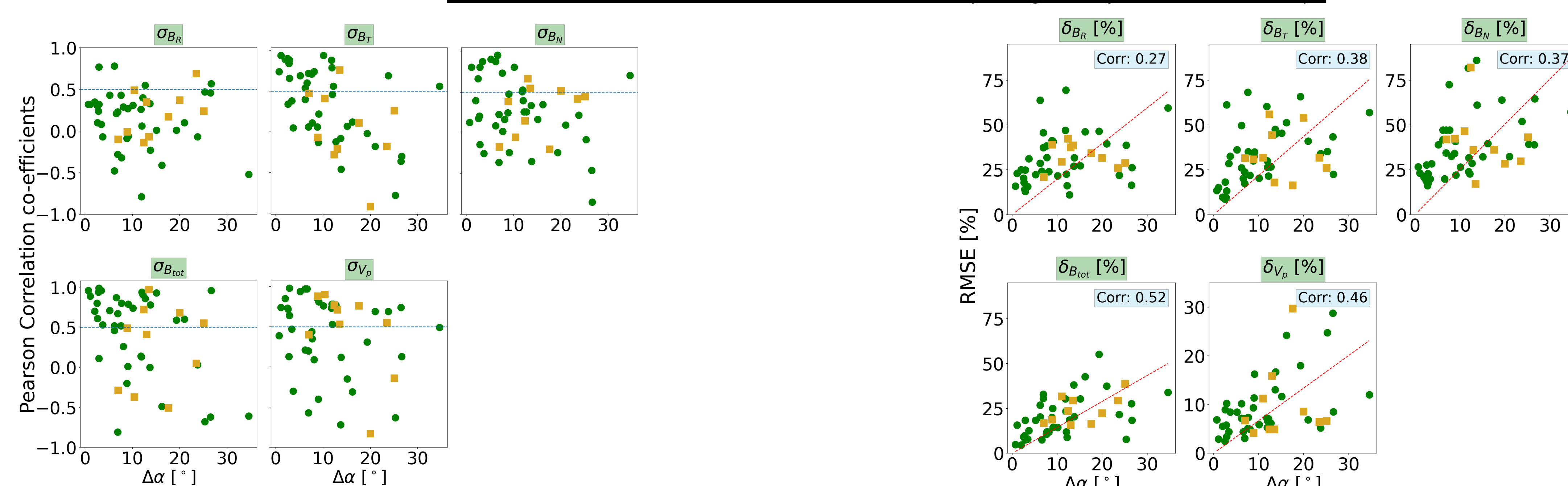
### Magnetic ejecta properties



- Average speed and mean magnetic field strength are consistent from L1 and STEREO-A.
- However, the expansion signatures such as expansion speed, Distortion parameter (DiP), and front-to-back ratio of magnetic field are low to moderately correlated at two points, highest for front-to-back ratio.

## 5. Analyzing Time-evolution

### Pearson correlation & normalized RMSE of magnetic field and velocity



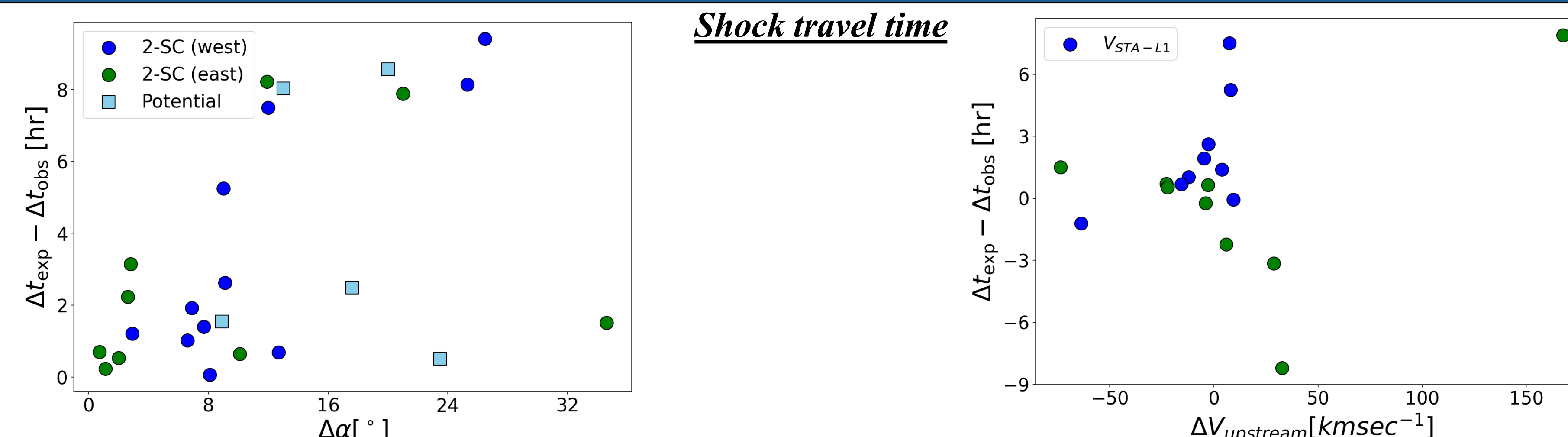
- Pearson correlation coefficient ( $\sigma$ ) between L1 and STEREO-A measurements decreases with increasing angular separation ( $\Delta\alpha$ ) with very large variability.
- Correlation for  $B_{tot}$  and  $V_p$  are highest; low or uncorrelated events could be due to local compression.
- Field components are not well correlated ( $<0.5$ ) even within  $10^\circ$  maybe due to small values.

- The normalized root mean square error (RMSE),  $\delta$ , between the L1 and STEREO-A measurements increases monotonically as the angular separation,  $\Delta\alpha$ .
- Potential events follow similar trend like 2-SC events.
- As beyond  $\sim 5^\circ$ , RMSE tends to be larger, more datapoints lie above the trend line at larger separation.

RMSE is more consistent than correlation for angular separation range in our study!

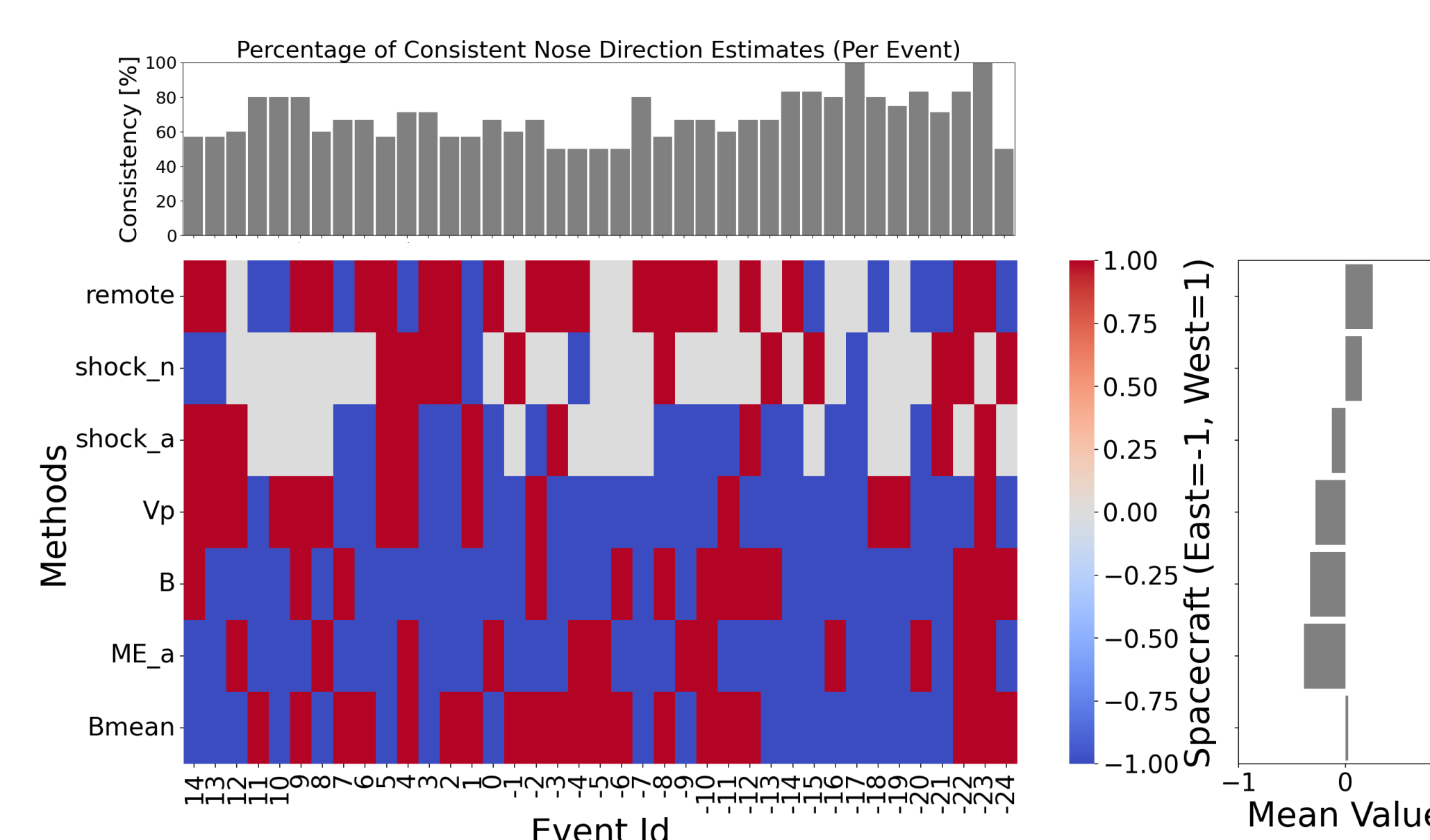
## 6. Shocks

### Shock travel time



- Among the 49 events (2-SC + Potential), 30 show a shock or shock-like disturbance at both spacecraft.
- The results indicate that shocks propagate largely radially outward, meaning STEREO-A can forecast shock arrival at Earth.
- Together, the figures suggest that shock behavior depends more on local conditions—such as upstream solar wind speed—than on angular separation.

### CME nose estimate using various methods



- Using 49 two-spacecraft CMEs, we estimated whether the CME nose lay to the east or west.
- We applied seven methods, spanning remote-sensing and in-situ approaches, including shock arrival and normal direction; ejecta arrival, proton speed, and mean/max magnetic field.
- A westward bias is observed in remote-sensing estimates while *in situ* methods predominantly indicate eastward directions.
- The consistency histogram indicate all seven methods agree for only two events.